### State-of-the-art modeling of impurity transport in stellarators

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- U. Oxford M. Barnes and F. Parra.
- Max Planck IPP H. Smith, R. Kleiber, J. Alcusón, A. Langenberg, P. Xanthopoulos, A. Zocco.
- NIFS S. Satake and K. Fujita.
- and W7-X team.

## Experimental motivation

- Stellarators count with a number of situations with reduced impurity core confinement that **standard neoclassical modeling** have not successfully modeled.
- These situations are best exemplified by the LHD impurity hole plasmas (left figure), the high-density-H (HDH) mode W7-AS plasma (right fig. ◊).



Standard neoclassical modelling:

- Radially local monoenergetic (r and v are mere parameters) DKES [Hirshman PF'86] based calculations.
- Trucated electric field (only radial,  $E_r$ ) and incompressible form of  $E_r \times B$  drift.
- Pitch-angle-scattering collisions, with momentum conservation schemes, of limited application to impurity calculations.

[Ida PoP'11] (left) ; [McCormick PRL'02] (right)

⇒ Recent intense code development and analytical effort to improve the modeling of collisional impurity transport.



2 Improvements on code development and analytical theory

3 Comparison with the experiment

4 Gyrokinetic quasilinear and nonlinear transport of impurities

5 Summary



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### Basic features of neoclassical electric fields in stellarators

Neoclassical radial fluxes (Γ) are not intrinsically ambipolar in stellarators for E<sub>r</sub> = 0. E<sub>r</sub> is constrained by:

$$\sum_{lpha} Z_{lpha} e \Gamma_{lpha}(E_r) = 0.$$

Over the flux surfaces unbalanced charge density leads the plasma to generate a potential φ<sub>1</sub> to preserve quasineutrality.

 $\varphi_1$  scaling at low  $\nu_{*i}$  [Calvo JPP'18]

- In stellarators φ<sub>1</sub> ~ ν<sub>\*i</sub><sup>-1</sup> in the 1/ν regime up to constant values reached at either at the √ν-regime (large E<sub>r</sub>) or superbanana-plateau regime (low E<sub>r</sub>).
- In tokamaks  $\varphi_1 \sim \nu_{*i}$  at low collisionality.
- The *E* × *B* and magnetic drifts become comparable as *Z* increases:

$$\mathbf{v}_m \cdot \nabla r \sim \mathbf{v}_{E1} \cdot \nabla r.$$

• The electrostatic and magnetic trapping become comparable as *Z* increases:

$$\mu \nabla_{\parallel} B \sim Z e \nabla_{\parallel} \Phi_1.$$



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### The impact $\varphi_1$ and its importance for impurity transport

- Impact of  $\varphi_1$  on  $\Gamma_Z$  confirmed for  $e\varphi_1/T_i \lesssim 0.05$  and low-Z impurities like C<sup>6+</sup> [Regaña PPCF'13, Regaña NF'17].
- The *phase* of  $f_{1Z}$  respect to  $(\mathbf{v}_m + \mathbf{v}_{E1}) \cdot \nabla r$  underlies the impact of  $\varphi_1$  on the  $\Gamma_Z$ .





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 $\Gamma_{Z} = \left\langle \int f_{1Z}(\mathbf{v}_{m} + \mathbf{v}_{E1}) \cdot \nabla r \mathrm{d}^{3} v \right\rangle$ 

• The *phase* of  $f_{1Z}$  respect to  $(\mathbf{v}_m + \mathbf{v}_{E1}) \cdot \nabla r$  underlies the impact of  $\varphi_1$  on the  $\Gamma_Z$ .



- Without  $\varphi_1$ .
- $(\mathbf{v}_m + \mathbf{v}_{E1}) \cdot \nabla r$  and  $n_{1Z}$  for  $C^{6+}$
- Radial (inward or outward) drifts do not preferentially overlap high or low C<sup>6+</sup> density regions.
- $\Rightarrow$  Weakier inward flux

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# What to improve respect to first EUTERPE simulations of $\varphi_1$ and $\Gamma_Z$ [Regaña NF'17]?

- EUTERPE is a  $\delta f$  PIC code, radially global on its gyrokinetic version.
- In this presentation we use the radially local <u>neoclassical version</u>.
- It solves

$$\frac{\partial f_1}{\partial t} + \mathbf{R}^{\mathbf{0}} \cdot \nabla f_1 + \dot{v}_{\parallel}^0 \frac{\partial f_1}{\partial v_{\parallel}} = -f_{\mathsf{M}} \left( \mathbf{v}_m + \mathbf{v}_{E1} \right) \cdot \nabla r \left[ \frac{n'}{n} + \frac{Ze}{T} E_r + \left( \frac{mv^2}{2T} - \frac{3}{2} + \frac{Ze}{T} \Phi_1 \right) \frac{T'}{T} \right] + \mathcal{C}(f).$$

$$\Phi_1 = \frac{T_e}{e} \left( n_{0e} + n_{0i} \frac{T_e}{T_i} \right)^{-1} \left( n_{1i} - n_{1e} \right).$$

- **1** The guiding center trajectory  $\dot{\mathbf{R}}^0$  does not retain the **magnetic drift** (neither the component accross the surface nor the tangential component  $\mathbf{v}_m \cdot \nabla \alpha$ ).
- $\Rightarrow$  Neglecting  $\mathbf{v}_m \cdot \nabla \alpha \rightarrow$  unrealistic situation with vanishing drift over flux surface when  $E_r = 0$ .
- $\Rightarrow$  Unphysically large size of  $f_1$  and its moments (including  $\varphi_1$ ) in the vicinity of  $E_r = 0$ .
- 2 The collision operator used included only pitch-angle-scattering, with questionable application to moderate to high collisional impurities.
- 3 Integration timestep has to be reduced a factor of  $O(\sqrt{m_e/m_i})$  when kinetic electrons need to be considered.

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### The importance of $\mathbf{v}_m\cdot \nabla \boldsymbol{\alpha}$ and the code KNOSOS

### KNOSOS [Velasco (2019) arXiv:1908.11615]

• **Bounce-averaged** local neoclassical code that, following [Calvo, PPCF'17], consistently solves at low collisionality and arbitrary geometry:

**Drift kinetic equation** (DKE)  $\rightarrow$  *g*<sup>*b*</sup> and fluxes like  $\Gamma$  and *Q*;

Ambipolarity condition  $\rightarrow E_r$ ;

**Quasi-neutrality** equation  $\rightarrow \varphi_1$ ;

- It includes analytically exact tangential magnetic drift  $\mathbf{v}_m \cdot \nabla \alpha$ .
- O(1) min. of computing time per flux surface,  $\times 2$  if kinetic electrons are included.



#### Effect of the $\mathbf{v}_m \cdot \nabla \alpha$ on $\varphi_1$ and benchmark for an LHD case

 $g_b$  is the departure of distribution function from a Maxwellian;  $\alpha = \theta - \iota \zeta$  labels magnetic field lines in the second structure of the second st



Alternatively, the full drift trajectories can be integrated, i.e. perform a radially global simulation.
 FORTEC-3D [Satake PFR'06], a PIC δf radially global (5D) code, has found corrections to EUTERPE's φ<sub>1</sub> as those found by KNOSOS in the situations where v<sub>m</sub> · ∇α can be important [Fujita PFR'19].



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### The importance of the collision operator and the code SFINCS



 The linearized impurity-main ion collision can be 
 √m<sub>i</sub>/m<sub>Z</sub>-expanded to see that three main pieces controls Γ<sub>Z</sub>: I) pitch-angle and energy diffusion; II) friction against main ions; III) main ion pressure anisotropy [Mollén PoP'15].

$$C_{Zi}^{l}[F_{1Z};h_{i}] = \underbrace{\frac{\nu_{Zi}T}{m_{Z}}\nabla_{\mathbf{v}}\cdot\left[F_{MZ}\nabla_{\mathbf{v}}\left(\frac{F_{Z1}}{F_{MZ}}\right)\right]}_{(III)} + \underbrace{\frac{\nu_{Zi}m_{Z}}{T}Av_{\parallel}}_{(III)} - \underbrace{\frac{\nu_{Zi}m_{Z}}{T}\left(\operatorname{tr}(\overset{\leftrightarrow}{M}) - \frac{m_{Z}}{T}\overset{\leftrightarrow}{M}:\mathbf{vv}\right)}_{(III)}.$$

- (II) and (III) depend on main ion distribution  $h_i$  through A and M.
- (II) is responsible for parallel momentum conservation; (III) can be important in stellarators [Calvo arXiv: 1907.08482, see **P1-05**].

### SFINCS [Landremann PoP'14]

- Local neoclassical (4D)  $\delta f$  continuum code that consistently solves, for arbitrary geometry, DKE, ambipolar condition and quasi-neutrality.
- It includes all these effects, as it implements the **linearized Fokker-Planck collision operator**.
- Multispecies at non-trace concentration.



EUTERPE PAS and mass-ratio-expanded collision operator (w/o III) vs. SFINCS .



- When the main ions are low collisional and the impurities are in the Pfirsch-Schlüter regime:
- $\rightarrow$  if  $\varphi_1 = 0$  the dependence of  $\Gamma_Z$  on  $E_r$  vanishes. [Helander PRL'17]  $\rightarrow$  more likely screening.
- → if  $\varphi_1 \neq 0$  the dependence of  $\Gamma_Z$  on  $E_r$  is recovered, even at low  $\varphi_1$  values [Calvo NF'18, Buller JPP'18].



$$\begin{split} &\Delta\varphi_1 = \varphi_1^{\max} - \varphi_1^{\min} \text{ for W7-X, } r/a = 0.8, \ T_i = 1.9 \text{ keV,} \\ &d_r \log T_i = -6.8, \ T_e = 1.3 \text{ keV, } d_r \log T_e = -6.5, \\ &n_i = 1.15 \times 10^{20} \text{ m}^{-3} \text{ } d_r \log n_i = -5.8. \end{split}$$

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- w/o φ<sub>1</sub>: The monotonic flux dependence on E<sub>r</sub> weakens, or equivalently Γ<sub>Z</sub>(E<sub>r</sub>) flattens.
- w/ φ<sub>1</sub>: Monotonic linear dependence Γ<sub>Z</sub>(E<sub>r</sub>) vanishes resulting on regions with weaker and enhanced inward flux.
- The numerically simulated flux level agree reasonably well with analytical prediction.
- To what extent moderation of inward (or even outward) Γ<sub>Z</sub> are expected when E<sub>r</sub> < 0? →</li>
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### Analytical formulas + KNOSOS

- Analytical formulas for Γ<sub>Z</sub> for trace impurities and large-aspect ratio in the mixed collisionality regime available in [Calvo NF'18, Buller JPP'19].
- Extended to include main-ion pressure anisotropy (III) for low  $\nu_{*i}$  and different collisionality regimes for impurities in [Calvo arXiv: 1907.08482] (see **P1-05**).
- Formulas depend on the main ion distribution *h<sub>i</sub>*, which need to be numerically worked out:

KNOSOS  $\xrightarrow{h_i}$  Analytical formulas for  $\Gamma_Z$ 

- In W7-X (top), weaker  $\varphi_1$  leads to  $V_Z > 0$  for  $\partial \log T_i/2 < \partial_r \log n_i$ .
- In LHD (bottom), stronger φ<sub>1</sub> leads to V<sub>Z</sub> > 0 restricted to ∂<sub>r</sub> log n<sub>i</sub> > 0.



 $V_{W^{44}}$  as function of  $\partial_r \log n_i$  and  $\partial_r \log T_i$ .  $n_e = 6.3 \times 10^{19} \text{ m}^{-3}$ ,  $T_b = 1.7 \text{ keV}$ . 600 runs. 2 Improvements on code development and analytical theory

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## LHD impurity hole with SFINCS

- (Q)
- The impurity hole plasma simulated with SFINCS [Mollén PPCF'18] (exact F-P collision operator, multispecies (at non-trace concentration), including  $\varphi_1$  yields:
- Huge corrections to standard neoclassical calculations.
- Impact goes in the opposite direction to what experiment indicates (inward Γ<sub>Z</sub> considerably enhanced).



F-P w/o  $\varphi_1$ ; PAS w/o  $\varphi_1$ ; F-P w/  $\varphi_1$ ; PAS w/  $\varphi_1$ ; DKES + momentum conserv.





- Past experiments have measured differences on plasma potential attributable to  $\Delta \varphi_1$  [Pedrosa NF'15].
- Doppler reflectometry (DR) has measured at different poloidal and toroidal positions appreciable difference on the total radial electric field [Estrada NF'19].
- Contribution from  $-d\Phi_1/dr$  to the total radial electric field calculated with EUTERPE [Regaña PPCF'18] and KNOSOS .
- Radial electric field contribution from -d<sub>r</sub>φ<sub>1</sub> is negligible with adiabatic electrons.
- Including kinetic electrons matters and -d<sub>r</sub>φ<sub>1</sub> correction become appreciable althought not to the level the DR system observes.





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## $\Gamma_Z$ in W7-X#180919.055



- Numerical analyses of W7-X discharges have recently been approached with these tools.
- The agreement between the codes is found where they are expected to agree.
- The differences between SFINCS and EUTERPE on  $\varphi_1$  or collision operator translate into small difference on neoclassical fluxes.



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- $D_{Ar^{16+}}$  is strongly underestimated. XICS-measured  $V_Z$  and neoclassical estimate coincide in sign, but disagree in a factor of O(10).
- Including classical flux, which can be important in W7-X [Buller JPP'19], strongly modifies  $V_Z$  in the mid outer regions  $\rightarrow$  to what extent the difference on  $V_Z$  and  $D_Z$  are attributtable to turbulence?

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### The code stella [1]

### Code characteristics

- Eulerian + semi-Lagrange, delta-f nonlinear gyrokinetic code.
- Mixed implicit-explicit scheme.
- Multiple ion species + kinetic electrons.
- General nested flux surface geometry (including interface with VMEC).
- Currently flux tube, full flux surface under development.
- Benchmarked against GS2. Ongoing stellarator specific benchmark against GENE.
- Coupled to SFINCS for neoclassical equilibrium and  $\varphi_1$  (coupling with KNOSOS foreseen).



Selected stella /GENE benchmark results: frequency  $\omega(k_x)$  for an ITG case w/ adiabatic e<sup>-</sup> (top); growth rate  $\gamma(k_y)$  for a TEM case (bottom) [A. G. Jerez *in progress*.]

[1] M. Barnes, F. I. Parra and M. Landreman J. Comput. Phys. (2019) の・モン・モン モー つへで

Impurity transport modeling



• In the trace impurity limit it can be written a transport law of the form:

$$\Gamma_Z^{\rm ql} = -n_Z \left( D_{Z1} \frac{\mathrm{d} \ln n_Z}{\mathrm{d} r} + D_{Z2} \frac{\mathrm{d} \ln T_Z}{\mathrm{d} r} + C_Z \right),$$

For sufficiently heavy impurities, with bounce (ω<sub>bZ</sub>) and drift (ω<sub>dZ</sub>) frequencies, in the presence of an instability with fluctuating potential φ, frequency ω and k<sub>⊥</sub>ρ<sub>i</sub> ≤ 1:

$$\omega/\omega_{bZ} \gg 1$$
 and  $\omega/\omega_{dZ} \gg 1$ .

• Then, the solution of the GK eq. for the impurity distribution function  $g_Z$  leads to:

$$D_{Z1} = \frac{\gamma k_{\alpha}^{2}}{\omega^{2} + \gamma^{2}} \left\langle \phi^{2} \Gamma_{0}(b) \right\rangle$$
$$D_{Z2} = \frac{\gamma k_{\alpha}^{2}}{\omega^{2} + \gamma^{2}} \left\langle \phi^{2} b \left[ \Gamma_{1}(b) - \Gamma_{0}(b) \right] \right\rangle$$
$$C_{Z} = 0$$

with  $b = k_{\perp}^2 T_Z / m_Z \Omega_Z^2$ ,  $\Gamma_n(b) = I_n(b)e^{-b}$ .

### Quasilinear transport of impurities with stella

- Accounting for the specific  $\phi$  structure, growth rate, frequency, magnetic geometry, parallel impurity dynamics (not included in analytical treatment) requires a numerical approach.
- Collisions are not included.

• W7-X (std. config.), r/a = 0.8 with  $T_i = T_e = 0.5$  keV,  $a/L_{T_i} = a/L_{T_e} = 6.0$  and  $a/L_{n_i} = a/L_{n_e} = 1.0$ .



• The calculated impurity radial flux of a linear GK simulation ( $\Gamma_Z$ ) allows, defining  $\Gamma_Z^{ql}$  as:

$$\Gamma_{Z}^{\rm ql} = \Gamma_{Z} \frac{\omega^{2} + \gamma^{2}}{\gamma k_{y}^{2} \langle \phi^{2} \rangle},$$

to work out the spectra, relative sizes and sign of  $D_{Z1}$ ,  $D_{Z2}$  and  $C_Z^{1}$ .

<sup>1</sup>For characterizing LHD impurity hole plasma see [Mikkelsen PoP'14].

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#### Impurity transport modeling

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- $D_{Z1} \Rightarrow$  **outward-directed**  $\Gamma_Z^{ql}$  while  $D_{Z2} \Rightarrow$  (much weaker) **inward-directed**  $\Gamma_Z^{ql}$ .
- Covection term  $C_Z$  introduces additional **inward-directed**  $\Gamma_Z^{ql}$  contribution.
- Weak mass or Z-dependence observed [Barnes PRL'12], with the exception of the  $D_{Z2}$  at low  $k_y$ .
- $\Gamma_Z^{ql}$  abruptly decreases once the electron-driven (ETG) instability is found at  $k_y \rho_i > 6$ .

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- Preliminary nonlinear analyses have begun.
- W7-X (std. config.) at r/a = 0.8 with  $T_i = T_e = 0.5$  keV,  $a/L_{T_i} = 6.0$  and  $a/L_{n_i} = 1.0$ . adiabatic electrons:



	Quasilinear	Nonlinear			
$D_{Z1}=D_Z$	> 0	1.6			
$D_{Z2}$	< 0	-0.02			
$C_Z$	> 0	-1.38			
V	< 0	1.1			
$\overline{[D_Z]} = \Gamma_{gB} n_i^{-1} m ; [C_Z] = [V] = \Gamma_{gB} n_i^{-1}.$					

- Nonlinear simulations show (and QL analyses capture too) strong outward-directed n'<sub>Z</sub>-driven diffusion, and weak inward-directed thermo-diffusion.
- Nonlinearly **anti-pinch** (*C<sub>Z</sub>*) is found, which contradicts QL analysis.

### Summary

• The basic understanding of collisional impurity radial transport has improved due to an intense code and analytical theory development in recent years.

	EUTERPE <sup>2</sup>	SFINCS	KNOSOS + analyt. $\Gamma_Z$	FORTEC-3D
Kind	Monte-Carlo $\delta f$	Eulerian $\delta f$	Bounce-averaged $\delta f_i$	Monte-carlo $\delta f_i$
Geometry	General/VMEC	General/VMEC	General/VMEC	General/VMEC
Local/global	Local	Local	Local	Global
$\mathbf{v}_{m}\cdot  abla lpha^{\dagger}$	None	Model	Analytically exact.	Included
C(f) for $Z$	$\sqrt{m_i/m_Z}$ -expanded <sup>3</sup>	Linearized Fokker-Planck	Analy. eval. Γ <sub>Ζ</sub>	Model for $C_{ab}$
C(f) for i,e	Model for $C_{aa}$	Exact Fokker-Planck	PAS <sup>4</sup>	Model for $C_{ab}$
CPU usage	Intense	Moderate/intense	Light	Intense

• The benchmark work has quantitatively confirmed the impact of neoclassical effects, traditionally neglected, that are important for an accurate evaluation of neoclassical  $\Gamma_Z$ .

- Experimental confirmation of predictions has been moderate (e.g.  $\varphi_1$  measurements, of convection  $V_Z$  in selected LHD plasmas) but also made evident big gaps, for instance,  $D_Z$ .
- First quasilinear and nonliner simulations carried out in order to study the role of gyrokinetic turbulence on  $\Gamma_Z$ .

- $^{3}w/o$  pressure anisotropy terms.
- <sup>4</sup>Pitch angle scattering.

<sup>&</sup>lt;sup>2</sup>Neoclassical version of the code EUTERPE

<sup>&</sup>lt;sup>†</sup>Tangential magnetic drift. Necessary for  $E_r \ll T/(ea)$ , when superbanana plateau regime is accessed [Calvo PPCF'17].

### Backup slides



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Quasilinear transport of impurities with stella (V and D).

• Expressing the flux in terms of the conventional diffusion and convection coefficient  $D_Z$  and  $V_Z$ :

$$\frac{\Gamma_Z^{\rm ql}}{n_Z} = -D_Z \frac{{\rm d}\ln n_Z}{{\rm d}r} + V_Z$$

 $V_Z$  can be obtained for a specific value of  $T'_Z$ . Setting  $a/L_{T_Z} = a/L_{T_i} = -6.0$  yields:









• Expressing the flux in terms of the conventional diffusion and convection coefficient  $D_Z$  and  $V_Z$ :

$$\frac{\Gamma_Z^{\rm ql}}{n_Z} = -D_Z \frac{{\rm d}\ln n_Z}{{\rm d}r} + V_Z$$

 $V_Z$  can be obtained for a specific value of  $T'_Z$ . Setting  $a/L_{T_Z} = a/L_{T_i} = -6.0$  yields:



•  $V_Z$  and  $D_Z$  have opposite sign, which would lead to peaked  $n_Z$  profiles.

•  $V_Z$  and  $D_Z$  have comparable size, which conflicts with the conclusion about  $V_Z$  and  $D_Z$  set by different mechanisms (NC or turb.) given their calculated values.